Adjustments to the law of the wall above an Amazonian Forest explained by a spectral link

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*Presentation for the Warm Boundary Layer Processes working group for the ASR program ARM/ASR - PI meeting, Washington DC, October 27, 2022 (email: gaby@duke.edu).

Acknowledgement: Department of Energy, Office of Science (DE-SC0022072)

Introduction

 The significance of the roughness sublayer (RSL) to a plethora of physical, chemical, and biological processes is not in dispute.



Figure from: https://doi.org/10.1016/j.ijft.2021.100077



Introduction

Focus here on **RSLABOVE CANOPIES** – where *multiple eddy types* dominate biosphere-atmosphere exchange



Roughness sublayer correction

• Correction **increases** mean velocity U relative to its log-law extrapolation



Roughness sublayer

• Correction to the 'law of the wall': $\frac{dU}{dz} \frac{\kappa(z-d)}{u_*} = \phi_{RSL} \left(\frac{z}{h}, \dots\right)$

- Log-law recovered when $\phi_{RSL}(.) = 1$
- Current approach:

Empirical coefficient

$$\phi_{RSL}(.) = 1 - exp \left| - \right|$$

$$\left[-\frac{a}{z^*}\right]$$

Thickness of the RSL



Problem:

• Many ϕ_{RSL} models proposed - but miss the key mechanism they purported to representenergetics of eddies.

Objective:

• Derive ϕ_{RSL} from energetics of turbulent eddies and compare with experiments in non-ideal settings (e.g. Amazonia).

Model Co-spectra from simplified budget

Eddy sizes or scales considered using the co-spectrum:



Two-term co-spectral budget

<u>Simplify</u>: High Reynolds number, stationary and equilibrium co-spectral budgets (i.e. turbulent transfer terms ignored).

Production of covariance $F_{wu}(k)$ at scale k:

 $\frac{dU}{dz} F_{ww}(k) +$

$$-\pi_u(k)=0$$

De-correlation due to pressure - velocity interaction (**requires closure at k**)

A Rotta type closure for $\pi_u(k)$

Models the *universal* tendency of <u>all</u> turbulent flows to return to isotropy at small scales

> Classical Rotta term **SLOW PART**

Isotropization of the production **FAST PART**



$$\pi_u(k) = -A_u \frac{F_{wu}(k)}{\tau(k)} - C_{Iu} \left| \frac{dU}{dz} F_{ww}(k) \right| \leq$$

 $\tau(k)$ = Onsager (1948) relaxation time scale

 A_u = Rotta constant (~1.8).

 C_{Iu} = Coefficient related to isotropization of production

<u>Rapid Distortion Theory</u> predicts $C_{Iu} = 3/5$ (Pope, 2000).



Production by

mean gradients

Stationary and equilibrium solution to the two-term co-spectral budget

This simplified budget relates the **co-spectrum** to the vertical velocity spectrum $F_{ww}(k)$.

$$F_{wu}(k) = \frac{1 - C_{IU}}{A_u} \left[\frac{dU}{dz} F_{ww}(k) \right] \tau(k)$$

 $\tau(k)$ = Onsager's (1948) relaxation time scale = $k^{-2/3} \epsilon^{-1/3}$

 $F_{ww}(k)$ = Vertical velocity energy spectrum, must be externally supplied (measured or assumed)

Shape of F_{ww}(k)

For canonical boundary layers (in the inertial layer):



The integrated co-spectral budget

$$\frac{dU}{dz} = \left(\frac{4}{7} \frac{1}{C_o} \frac{A_u}{1 - C_{IU}}\right)^{3/4} \frac{\overline{-(w'u')}}{z}^{1/2},$$



Upon integration yields the **Prandtl-von Karman** velocity profile:

$$U(z) = \left(\frac{4}{7}\frac{1}{C_o}\frac{A_u}{1 - C_{IU}}\right)^{\frac{3}{4}}u_*\log(z) + B$$

The von Karman constant can be estimated

$$\kappa = \left(\frac{4}{7} \frac{1}{C_o} \frac{A_u}{1 - C_{IU}}\right)^{-\frac{3}{4}} \approx 0.36$$
(24/55)(1.5) 3/5

1 2

Estimating RSL correction

DETAILED MODEL - Measured $F_{ww}(k)$

$$\phi_{RSL}(.) = \left(-\frac{\overline{u'w'}}{{u_*}^2}\right) \left(\frac{A_u}{1-C_{IU}}\right) \frac{u_*\kappa(z-d)}{\int_0^\infty \tau(k) F_{ww}(k)dk}$$

SIMPLIFIED MODEL:

$$\phi_{RSL}(.) = \left(-\frac{\overline{u'w'}}{{u_*}^2}\right) \left(\frac{A_u}{1-C_{IU}}\right) \frac{u_*\kappa(z-d)}{\tau_{eff}\,\sigma_w^2}; \quad \tau_{eff} = \frac{2\sigma_w^2}{\epsilon(z)}$$

Emergence of a 'macro-dissipation' length in $\phi_{RSL}(.)$

$$\phi_{RSL}(.) = \frac{1}{2} \left(-\frac{\overline{u'w'}}{{u_*}^2} \right) \left(\frac{A_u}{1 - C_{IU}} \right) \left(\frac{u_*}{\sigma_w} \right)^4 \frac{L_{BL}}{L_d};$$

$$BL = \kappa (z - d); \ L_d = \frac{{u^*}^3}{\epsilon(z)}.$$

This is a new scale for RSL that differs from the 'canonical' shear length scale derived from *mixing layer* analogy (L_s) put forth by Raupach, Finnigan and others.



$$L_s = \frac{U}{dU/dz}$$





Experiments

23 March 2014 to 16 January 2015

GOAmazon (K34)

25 October to 25 November of 2015

ΑΤΤΟ



Amazon Tall Tower Observatory

At both sites: h=35 m and LAI = 6

GoAmazon: From Fuentes et al. (2016) ATTO: https://commons.wikimedia.org/wiki/File:Amazon_Tall_Tower_Observatory.jpg



kz

Breakpoints in spectra (ATTO)

ATTO

10¹

kL_S



kL_d

 10^{4}

Model comparisons



Conclusions

 An eddy viscosity that accommodates energetics of turbulence - analogous to the fluctuationdissipation theorem in statistical mechanics

$$v_t = \left(\frac{1 - C_{IU}}{A_u}\right) \int_0^\infty \tau(k) \ F_{ww}(k) dk.$$

- Emergence of a macro-scale dissipation length (L_d) that explains transitions in $F_{ww}(k)$ as well as RSL correction functions (ϕ_{RSL}) .
- Derived ϕ_{RSL} appears robust to non-ideal conditions at the two forested sites in Amazonia.
- <u>Future work</u>: include thermal stratification



EXTRA SLIDES

Integration of co-spectrum across all k

$$F_{wu}(k) = \frac{1 - C_{IU}}{A_u} \left[\frac{dU}{dz} F_{ww}(k) \right] \tau(k)$$

$$\overline{w'u'} = \int_0^\infty F_{wu}(k)dk = \int_0^{k_a} F_{wu}(k)dk \qquad +$$

Integration limit applicable for $Re_* \rightarrow \infty, \eta \rightarrow 0$ $F_{wu}(k)dk$ k_a

Large scales (attached)

 $F_{WW}(k) = C_o \epsilon^{2/3} k_o^{-5/3}$

$$F_{WW}(k) = C_o \epsilon^{2/3} k^{-5/3}$$

Assume turbulent kinetic energy (TKE) budget is in equilibrium so that

$$\epsilon = -\overline{w'u'}\frac{dU}{dz}$$

Dissipation of TKE = Production of TKE



Solution for the inertial subrange

-1/3

$$F_{ww}(k) = C_o \epsilon^{2/3} k^{-5/3}$$

$$\tau(k) = k^{-2/3} \epsilon^{2/3} k^{-5/3}$$

$$F_{wu}(k) = \frac{1 - C_{IU}}{A_u} \left[\frac{dU}{dz} F_{ww}(k) \right] \tau(k)$$

$$F_{wu}(k) = \left(C_o \frac{1 - C_{IU}}{A_u} \right) \left[\frac{dU}{dz} \right] \epsilon^{1/3} k^{-7/3}$$

NOTES:

If spectrum of vertical velocity scales as -5/3, then the co-spectrum scales as -7/3 – consistent with experiments and Lumley's (1967) arguments.

Also suggestive that turbulent transfer term $T_{wu}(k)$ may be less important if a -7/3 power-law prevails in co-spectrum.

Comparison to Lumley (1967)

$$F_{wu}(k) = C_{uw} \left[\frac{dU}{dz} \right] \epsilon^{1/3} k^{-7/3}$$
$$F_{wu}(k) = \left(C_o \frac{1 - C_{IU}}{A_u} \right) \left[\frac{dU}{dz} \right] \epsilon^{1/3} k^{-7/3}$$

Lumley's result Dimensional analysis

Co-spectral budget

Accepted
range
$$C_{uw} = 0.15 - 0.16 = \left(\frac{1 - C_{IU}}{A_u}C_o\right) = 0.145$$

I.8

This result establishes a link between the similarity constant in the *Lumley's co-spectrum* and the *'collage'* of well-established constants in turbulence.

Support for the -7/3 co-spectral exponent

Field Experiments (Kansas)

Quart. J. R. Met. Soc. (1972), 98, pp. 590-603

Lab Experiments (NASA)

333

551.510.522 : 551.551.8 J. Fluid Mech. (1994), vol. 268, pp. 333-372 Copyright © 1994 Cambridge University Press

Cospectral similarity in the atmospheric surface layer

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Local isotropy in turbulent boundary layers at high Reynolds number

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